

## On NS5-brane instantons and volume stabilization

---

**Hugo Looyestijn and Stefan Vandoren**

*Institute for Theoretical Physics and Spinoza Institute,  
Utrecht University, 3508 TD Utrecht, The Netherlands  
E-mail: [h.t.looyestijn](mailto:h.t.looyestijn), [s.vandoren@phys.uu.nl](mailto:s.vandoren@phys.uu.nl)*

**ABSTRACT:** We study general aspects of NS5-brane instantons in relation to the stabilization of the volume modulus in Calabi-Yau compactifications of type II strings with fluxes, and their orientifold versions. These instantons correct the Kähler potential and generically yield significant contributions to the scalar potential at intermediate values of string coupling constant and volume. Under suitable conditions they yield uplifting terms that allow for meta-stable de Sitter vacua.

**KEYWORDS:** Superstring Vacua, Flux compactifications.

---

## Contents

<b>1. Introduction</b>	<b>1</b>
<b>2. Volume stabilization</b>	<b>3</b>
2.1 Fluxes and D3-brane instantons	5
2.2 Adding NS5-brane instantons	6
2.3 Analysis of the model	8
2.4 Fivebranes and orientifold projections	10
<b>3. The N=2 scenario</b>	<b>11</b>
3.1 Gauged and ungauged N=2 supergravity	12
3.2 Including NS5-brane corrections	13
3.3 Truncation to $N = 1$	14
<b>A. Calculations in <math>N = 1</math></b>	<b>16</b>
<b>B. Potentials in <math>N = 2</math></b>	<b>19</b>
<b>C. The Przanowski metric</b>	<b>20</b>
C.1 Solutions to the master equation	20
C.2 Moment maps	21

---

## 1. Introduction

Flux compactifications and moduli stabilization in string theory has been an active area of research in recent years. For some reviews on the topic, see [1–3]. Of particular interest is the stabilization of the volume modulus of the compactification manifold. In the effective supergravity approximation, one assumes that curvatures are small and hence the volume of the internal space is large compared to the scale set by the string length  $l_s = 2\pi\sqrt{\alpha'}$ . Stabilization by finding the minima of the scalar potential of the low-energy effective action must therefore lead to a large value for the volume for this approach to make sense. This is one of the basic assumptions in the KKLT scenario [4], or the more recent so-called “large volume scenarios” (LVS) [5, 6]. The scale set by the volume has important consequences for low-energy physics, such as supersymmetry breaking and inflation, see e.g. [6–10].

For type IIB flux compactifications on Calabi-Yau (CY) manifolds (and their orientifolds), the volume is stabilized by nonperturbative effects, such as stringy D3-brane instantons whose Euclidean worldvolume wraps a four-cycle in the CY. Such an instanton stabilizes the volume of the corresponding four-cycle and therefore the associated Kähler

modulus. The relation between the four-cycle and two-cycle volumes is known in principle, but is complicated in practice since it requires inverting a set of coupled quadratic equations involving the triple intersection numbers, as we review in the next section. Therefore, although stabilization of all Kähler moduli indeed stabilizes the entire volume of the CY, it requires a case by case analysis to fix it at large values,<sup>1</sup> see for instance [9] for a recent analysis.

The fact that a wrapped Euclidean  $p$ -brane can stabilize the volume of a  $p + 1$ -cycle naturally raises the question of whether the overall CY-volume can be stabilized directly by wrapping a Euclidean fivebrane over the entire CY. From this point of view, one could expect NS5-brane instantons to play an important role in relation to volume stabilization.<sup>2</sup> In case such instantons contribute to the low-energy scalar potential, they will do so with exponentially suppressed terms of the form

$$\exp[-S_{\text{NS5}}] = \exp\left[-\frac{\mathcal{V}}{g_s^2}\right], \tag{1.1}$$

where  $S_{\text{NS5}}$  is the (real part of the) one-instanton action,  $\mathcal{V}$  is the volume of the CY in dimensionless units, and  $g_s$  is the ten-dimensional string coupling constant. This can be compared to the contribution of a D3-brane instanton, of the form

$$\exp[-S_{\text{D3}}] = \exp\left[-\frac{\text{vol}(\gamma_4)}{g_s}\right]. \tag{1.2}$$

Here,  $\text{vol}(\gamma_4)$  is the volume of the four-cycle  $\gamma_4$ .

At weak string coupling constant, one expects the D3-brane instantons to dominate over the NS5-brane instantons, such that one can safely ignore the latter. The two exponents only are of the same order of magnitude when

$$\frac{\text{vol}(\gamma_4)}{\mathcal{V}} = \frac{1}{g_s}. \tag{1.3}$$

For small four-cycles and weak string coupling,  $g_s \leq 1$ , this is never satisfied. However, there are plenty of CY manifolds which have four-cycles with larger volume than the total volume, as we review in the next section, so some care is needed to make this argument, especially when  $g_s \sim 1$ . Similar considerations hold for compactifications of type IIA strings on CY threefolds, in which membrane instantons arise by wrapping Euclidean D2-branes over three-cycles.

There is another reason to be careful in ignoring the fivebrane instantons. Assuming that both instantons contribute to the scalar potential in the effective action and  $\text{vol} \gamma_4 < \mathcal{V}$ , the exponents above can still be multiplied by prefactors to make them of the same order,

---

<sup>1</sup>The situation might seem to look better in type IIA theories, since there one can stabilize the Kähler moduli directly at the classical level by switching on fluxes [11]. However, the Kähler moduli are fixed again by solving a set of quadratic equations involving the triple intersection numbers, see equation (4.36) in [11]. So to find large values for the total volume, one ends up with similar difficulties as in type IIB.

<sup>2</sup>In  $N = 2$  compactifications of type IIB, one also expect D5-brane instantons to contribute. However, after orientifold projection with O3/O7 planes, the D5 brane is not BPS and therefore harder to analyze. Furthermore, we look for a mechanism that also applies to IIA and heterotic string theories.

especially at intermediate string coupling  $g_s \sim 1$ . We will show this more explicitly in the next section, for values  $\mathcal{V} \approx 100$  (which are the typical values of the original KKLT approach). The existing vacua of the LVS scenarios at  $\mathcal{V} \approx 10^{13}$  are not effected by NS5-brane instantons. However, at smaller volumes  $\mathcal{V} \sim 100$ , additional vacua can arise with interesting properties.

The purpose of this paper is to analyze the effects of NS5-brane instantons in relation to the stabilization of the volume modulus. In particular, we show that under certain conditions, the contributions from NS5-brane instantons yield uplifting terms in the scalar potential that can lead to meta-stable de Sitter vacua. Most of the work on moduli stabilization has focused on  $N = 1$  supersymmetry in four dimensions, as they give rise to semi-realistic string vacua. In such models, like e.g. type IIB strings on Calabi-Yau orientifolds, moduli can be stabilized by combining the effects of fluxes and quantum corrections coming from perturbative corrections to the Kähler potential and D3-brane instanton corrections to the superpotential. However, the Kähler potential is subject to higher loop corrections in  $\alpha'$  and  $g_s$  which are not known explicitly. For a recent discussion on this, see [9, 12, 13]. As we will show, the nonperturbative corrections of the form (1.1) also contribute to the Kähler potential, and are generically subleading with respect to the first perturbative corrections, but could compete with next-to-leading perturbative corrections. For this reason, our investigations are more meaningful in  $N = 2$  models, since in that case higher order corrections are absent due to the constraints from  $N = 2$  supersymmetry. This fact also motivated the authors of [14] to study  $N = 2$  moduli potentials in type IIA flux compactifications. These toy models can serve as good approximations for the more realistic  $N = 1$  string vacua. Moreover, for  $N = 2$  theories in type IIA, there are some explicit results known about the contribution of NS5-brane instantons [15–18] to the effective action for the hypermultiplets.

The plan of the paper is as follows. In section 2, we present the generic form of an NS5-brane instanton correction to the scalar potential. We study this in the setting of  $N = 1$  supergravity in four dimensions, and investigate the relation with the KKLT and LVS scenarios. In section 3, we discuss IIA strings compactified on a (rigid) CY, for which there is some explicit knowledge on NS5-brane instantons. We investigate the stability of the volume, and find the possibility that NS5-brane instantons can produce de Sitter vacua. We then truncate this model preserving local  $N = 1$  supersymmetry, and determine the Kähler and superpotential.

## 2. Volume stabilization

In this section, we review certain aspects of the KKLT scenario and discuss some of the subtleties that can arise in stabilizing the volume at large values in IIB orientifold compactifications. We then include terms that mimic the contributions from NS5-brane instantons to the Kähler potential, and re-analyze the stabilization of the volume modulus.

We consider an orientifold of type IIB string theory on a CY 3-fold with O3/O7 planes. The cohomology groups  $H^{(p,q)}$  are split under the orientifold mapping into odd and even

forms, and hence their dimensions split as  $h^{p,q} = h_+^{p,q} + h_-^{p,q}$ . We follow the notation of [19], although we change a few names and numerical factors.

Let us first list the various chiral fields. The field  $\tau$  contains the axion and dilaton, and is defined by  $\tau = l + ie^{-\phi_{10}}$ . The fields  $T_i$  are defined by

$$T_i = \tau_i + ih_i - 2\zeta_i, \quad i, j = 1, \dots, h_+^{1,1}, \quad (2.1)$$

where

$$\zeta_j = -\frac{i}{2(\tau - \bar{\tau})} d_{jab} G^a (G - \bar{G})^b, \quad G^a = c^a - \tau b^a, \quad a, b = 1, \dots, h_-^{1,1}. \quad (2.2)$$

The  $\tau_i$  capture the sizes of the even four-cycles under the orientifold projection and  $h_i$  are real field that arise by expanding the  $C_4$  gauge field over these four-cycles. The fields  $b^a, c^a$  are the expansions of the  $B_2$  and  $C_2$  forms respectively over the  $h_-^{1,1}$  cycles. Notice that the definition of  $\zeta_i$  contains intersection numbers of even and odd two-cycles.

The four-cycles  $\tau_i$  are related to the two-cycles  $t^i$  by the triple intersection numbers  $d_{ijk}$  as

$$\tau_i = d_{ijk} t^j t^k. \quad (2.3)$$

The total volume  $\mathcal{V}$  is only implicitly known as a function of the  $N = 1$  chiral coordinates through the relation

$$\mathcal{V} = \frac{1}{6} d_{ijk} t^i t^j t^k.$$

To write the volume in terms of the chiral fields we first use the definitions (2.1), (2.2) to find

$$\tau_i = \frac{1}{2}(T_i + \bar{T}_i) - \frac{i}{2} d_{iab} b^a b^b (\tau - \bar{\tau})$$

or in terms of the chiral fields

$$\tau_i = \frac{1}{2}(T_i + \bar{T}_i) - \frac{i}{2} \frac{1}{(\tau - \bar{\tau})} d_{iab} (G - \bar{G})^a (G - \bar{G})^b.$$

One then solves the quadratic equations in equation (2.3) to obtain functions  $t^i(\tau_j)$ , and one obtains the volume  $\mathcal{V}$  depending on the chiral fields via  $\{\tau - \bar{\tau}, T_i + \bar{T}_i, (G - \bar{G})^a\}$ .

The type IIB Kähler potential is given by [20, 19]

$$K = K_{\text{cs}}(U, \bar{U}) + K_{\text{k}}(\tau, T, G) \quad (2.4)$$

$$K_{\text{k}} = -\ln[-i(\tau - \bar{\tau})] - 2 \ln \left[ \mathcal{V}(\tau, T, G) + \xi \text{Im}(\tau)^{3/2} \right]. \quad (2.5)$$

The Kähler potential  $K_{\text{k}}$  in (2.5) is the tree-level expression, together with the leading perturbative  $\alpha'$  correction proportional to the parameter  $\xi = -\frac{\chi(CY)\zeta(3)}{2(2\pi)^3}$ , containing the Euler number  $\chi(CY)$  of the internal Calabi-Yau  $M$  (we use conventions where  $l_s = 2\pi\sqrt{\alpha'}$ ). The complex structure deformations  $U$  are described by  $K_{\text{cs}}$ , whose precise form is not important. Higher string loop corrections could give a dilaton-dependence to  $K_{\text{cs}}(U, \bar{U})$ , but this is beyond the approximation we are working in.

The scalar potential for a Kähler potential  $K$  and superpotential  $W$  is given by

$$V = e^K \left( K^{\alpha\bar{\beta}} D_\alpha W D_{\bar{\beta}} \bar{W} - 3|W|^2 \right), \quad (2.6)$$

where the indices  $\alpha, \bar{\beta}$  run over all chiral fields, with  $K^{\alpha\bar{\beta}}$  the inverse Kähler metric.

## 2.1 Fluxes and D3-brane instantons

The tree-level superpotential is given by

$$W = W_0(\tau, U) = \int_M \Omega \wedge G_3, \tag{2.7}$$

where  $G_3$  is the complex combination of the background three-form fluxes, given by  $G_3 = F_3 - \tau H_3$ .

We now assume that the complex structure moduli  $U$  and the axio-dilaton  $\tau$  are stabilized at a higher energy scale at a SUSY minimum, by demanding  $D_\tau W = D_U W = 0$ . To stabilize the Kähler moduli we add the nonperturbative instanton corrections of wrapping Euclidean D3-branes over four-cycles. The superpotential is then given by [4].

$$W = W_0 + \sum_i A_i e^{-a_i T_i}, \tag{2.8}$$

where  $A_i$  and  $a_i$  are treated as field-independent parameters. In the literature one often considers the case where the  $G^a$  fields are absent. Using the expressions (2.4) and (2.8), one finds the following scalar potential [20, 21]

$$V = e^K \left[ K^{j\bar{k}} \left( a_j A_j a_k \bar{A}_k e^{-a_j T_j - a_k \bar{T}_k} - (a_j A_j e^{-a_j T_j} \bar{W} K_{\bar{k}} + \text{c.c.}) + K_j K_{\bar{k}} |W|^2 \right) - 3|W|^2 \right]. \tag{2.9}$$

where  $K^{i\bar{j}}$  are the components of the inverse metric  $K^{\alpha\bar{\beta}}$  in the directions of the Kähler moduli. In the absence of the  $G^a$  fields, there is no-scale structure at tree-level, leading to  $K^{i\bar{j}} K_i K_{\bar{j}} = 3$ . This no-scale structure is broken when  $\alpha'$ -corrections are included, and one then finds (see [20], or for some further details of the calculation, see appendix A)

$$V = e^K \left[ K^{j\bar{k}} \left( a_j A_j a_k \bar{A}_k e^{-a_j T_j - a_k \bar{T}_k} - (a_j A_j e^{-a_j T_j} \bar{W} K_{\bar{k}} + \text{c.c.}) \right) + 3\xi \frac{\xi^2 + 7\xi\mathcal{V} + \mathcal{V}^2}{(\mathcal{V} - \xi)(2\mathcal{V} + \xi)^2} |W|^2 \right]. \tag{2.10}$$

The various studies (see for example [22]) of these potential indicate a large volume AdS vacuum, which can be realized in explicit models. For example, in the  $\mathbb{P}^4_{[1,1,1,6,9]}$  model, which yields two Kähler moduli, the volume is expressed in terms of the 4-cycle volumes  $\tau_s, \tau_b$  as

$$\mathcal{V} = \frac{1}{9\sqrt{2}} \left( \tau_b^{3/2} - \tau_s^{3/2} \right). \tag{2.11}$$

This already gives an example where a four-cycle volume can be bigger than the total volume, as mentioned below (1.3).<sup>3</sup>

To remain in the regime of a geometrical compactification, one needs  $\tau_b > \tau_s$ . To obtain a large volume, one needs to arrange  $\tau_b \gg \tau_s$ . This is then solved self-consistently:

---

<sup>3</sup>Take for example the values  $\tau_s \sim 4.6$  and  $\tau_b \sim 120$  which give a total volume  $\mathcal{V} \sim 100$ .

one assumes a cycle to be small and the other large, approximates the potential in this regime and then searches for sets of vacua. In such a two-Kähler model this is doable, but a more general model will have many different Kähler moduli. To express the volume in terms of the four-cycles, one first has to solve the system of many coupled quadratic equations (2.3) and use the explicit form of the triple intersection numbers. Even for simple models with a few Kähler moduli, this can lead to equations which are not solvable analytically. Numerical methods can be used, but have their own limitations. One then has to find a limit on the four-cycles that leads to a large volume  $\mathcal{V}$ . This will be very difficult without any analytical control. Overall, it seems desirable to have a different mechanism that stabilizes the volume at once, without the need to stabilize the individual cycles that build up the total volume. A prime candidate for such a mechanism is the NS5-brane instanton, to which we turn now.

## 2.2 Adding NS5-brane instantons

We now add a correction due to the wrapping of an Euclidean NS5-brane over the entire CY. Such an instanton configurations contributes to correlators proportional to  $\exp(-\mathcal{V}/g_s^2)$ , where  $g_s$  is the 10-dimensional string coupling constant. The volume can not be expressed as a holomorphic function of the  $N = 1$  chiral fields. We therefore expect that the NS5-brane does not correct the superpotential, but instead it will correct the Kähler potential,

$$K_{\text{NS5}} = B \mathcal{V}^n \exp(-\mathcal{V}/g_s^2) = B \mathcal{V}^n \exp\left(\frac{1}{4}\mathcal{V}(\tau - \bar{\tau})^2\right). \quad (2.12)$$

The factor of  $\mathcal{V}^n$  represents the leading power of the instanton measure and the one-loop determinant of the fluctuations around the instanton solution. There is a proportionality factor  $B$  that could — in principle — depend on the moduli  $G^a$  and the dilaton  $\tau$ . We expect no dependence on the complex structure moduli  $U$  since the NS5-brane cannot probe the individual 3-cycles. Furthermore, before the orientifold projection, the NS5-branes correct the moduli space of Kähler deformations, and not the complex structure deformations. The prefactor  $\mathcal{V}$  is absent in an instanton corrected superpotential: because the superpotential is a holomorphic function of the chiral fields  $T_i$ , any non-trivial function  $A_i(T)$  in (2.8) breaks the shift symmetry on the imaginary parts  $h_i$  completely and is therefore forbidden. Since instantons are expected to break the shift symmetries to a discrete subgroup only, we can use superpotentials of the form  $\exp(-a_i T_i)$ . The Kähler potential does not need to be holomorphic, and we can therefore only use the real parts of the chiral fields  $T_i$ , which combine into powers of the volume  $\mathcal{V}$ .

A further argument in favor of (2.12) comes from the parent  $N = 2$  theory. NS5-branes correct the hypermultiplet moduli space even in the absence of fluxes [23]. In IIB compactifications, the hypermultiplet moduli are counted by the Kähler moduli, and the kinetic terms of these scalars receive corrections from NS5-brane instantons. After the orientifold projection, one expects these corrections to survive, even after fluxes are turned on. Our claim is that, to leading order, they enter the Kähler potential in they way described in (2.12). We elaborate further on this the next section for type IIA compacti-

fications, where we can determine the one-instanton NS5-brane contribution explicitly in some special cases [15].

We have to ask in which regime this approach makes sense. We want to remain in the one-instanton regime, and not consider multiple instanton contributions since nothing is known about them. Because an NS5-brane instanton scales as  $\exp(-\mathcal{V}/g_s^2)$ , this requires that

$$\frac{g_s^2}{\mathcal{V}} < 1. \tag{2.13}$$

However, since the NS5-brane is the magnetic dual of the fundamental string, one also has to consider perturbative effects, which we expect to organize an expansion in powers of  $g_s^2/\mathcal{V}$ . Besides that, there are also higher order  $\alpha'$  corrections which are left out. Only the first perturbative correction is included in our analysis in (2.4). In  $N = 2$  theories these higher string-loop corrections are absent [24, 25], which makes the discussion of the NS5-brane instanton more reliable. We discuss such models in the next section. We take a pragmatic approach here and isolate the NS5-brane instanton correction from all other corrections in the Kähler potential  $K_0$ . Hence we write

$$K = K_0(\tau, T, U, G) + K_{\text{NS5}}(\tau, T, G), \tag{2.14}$$

and expand to leading order in  $K_{\text{NS5}}$ . To obtain the expression for the inverse Kähler metric  $K^{\alpha\bar{\beta}}$  (where  $\alpha = \{\tau, T, G, U\}$  lists all the chiral fields) in the direction of the Kähler moduli, we use

$$K^{i\bar{j}} = K_0^{i\bar{j}} - K_{\text{NS5}}^{i\bar{j}}, \quad K_{\text{NS5}}^{i\bar{j}} \equiv K_0^{i\bar{\alpha}} K_{\text{NS5}\bar{\alpha}\beta} K_0^{\beta\bar{j}}, \tag{2.15}$$

and now  $K^{i\bar{j}} K_{\bar{j}k} = \delta_k^i + \mathcal{O}(K_{\text{NS5}}^2)$ .

In principle, there could be a dependence in  $K_{\text{NS5}}^{i\bar{j}}$  on the fields  $b^a$ , which enter through the definition of the chiral fields  $G^a$  in (2.1) and (2.2). As explained at the end of appendix A, the exact dependence on  $b^a$  is subleading in  $K_{\text{NS5}}^{i\bar{j}}$ .

To obtain the scalar potential, we work out equation (2.6), setting  $D_\tau W = D_U W = 0$ . For simplicity, we now set  $G^a$  to zero. It would be interesting to consider the effects of non-zero  $G^a$ , but this is beyond the scope of this article. Using this in (2.6) leads to

$$V = V_0 + V_0 K_{\text{NS5}} - e^{K_0} K_{\text{NS5}}^{i\bar{j}} |D_i^{(0)} W|^2 + e^{K_0} K_0^{i\bar{j}} \left( \partial_i K_{\text{NS5}} W D_{\bar{j}}^{(0)} \bar{W} + \text{c.c.} \right). \tag{2.16}$$

The potential in absence of  $K_{\text{NS5}}$  is denoted  $V_0$ ,  $D_i^{(0)} W = \partial_i W + (\partial_i K_0) W$  and  $K_0^{i\bar{j}}$  is the inverse of  $K_{0,\alpha\bar{\beta}}$  in the directions of the Kähler moduli. Remind that all chiral fields are labeled by  $\alpha, \bar{\beta}$ , and the Kähler moduli are a subsector thereof labeled by  $i, \bar{j}$ . Formula (2.16) in fact holds for any perturbation of the Kähler potential, labeled by  $K_{\text{NS5}}$ .

From equations (2.15)–(2.16) we can see that we have to compute  $K_{\text{NS5}\bar{\alpha}\beta}$ , so we have to take derivatives with respect to all chiral fields. This is to be contrasted to the situation in which one modifies the superpotential: the derivatives with respect to  $\tau$  and  $U$  are contained in  $D_\tau W$  and  $D_U W$ , which are set to zero.



Suppose for simplicity that  $W$  does not depend on the Kähler moduli (such as in equation (2.7)), then  $D_i^{(0)}W_0 = K_i^{(0)}W_0$  and hence

$$V = V_0 + V_0 K_{\text{NS5}} - e^{K_0} K_{\text{NS5}}^{i\bar{j}} |\partial_i K_0|^2 |W_0|^2 + e^{K_0} K_0^{i\bar{j}} \left( \partial_i K_{\text{NS5}} \partial_{\bar{j}} K_0 + \text{c.c.} \right) |W_0|^2. \quad (2.17)$$

Using expression (2.4) for  $K_0$ , we find that  $K_0^{i\bar{j}}$  is proportional to  $t^i t^{\bar{j}}$  and  $d^{i\bar{j}}$  (the inverse of  $d_{ij} = d_{ijk} t^k$ ), and both  $\partial_i K_0$  and  $\partial_i K_{\text{NS5}}$  are proportional to  $d_i = d_{ij} t^j$ . Upon contracting indices, this will combine nicely into powers of  $\mathcal{V}$ . The precise calculation can be found in appendix A. Ignoring the D3-brane instantons, and other subleading corrections to the Kähler potential, the leading correction to  $V_0$  is found to be

$$V_{\text{NS5}} = -\frac{9}{8} B |W_0|^2 g_s^{-3} \mathcal{V}^n \exp(-\mathcal{V}/g_s^2) \equiv \hat{B} \mathcal{V}^n \exp(-\mathcal{V}/g_s^2).$$

In the last expression, we have absorbed all the constants into a new prefactor  $\hat{B}$ .

### 2.3 Analysis of the model

We have motivated the scalar potential

$$V = V_0(\mathcal{V}) + \hat{B} \mathcal{V}^n e^{-\mathcal{V}/g_s^2}, \quad (2.18)$$

where  $\hat{B}, n$  are constants.

A general situation achieved in moduli stabilization is indicated on the left plot below, which displays an AdS minimum. We will investigate the physics of the NS5-brane contribution, depending on the signs of the parameters  $\hat{B}$  and  $n$ . The results are summarized in the table below. The most interesting case is when  $n$  is positive and  $\hat{B}$  is positive (the upper line in the middle graph). The exact location of the bump depends on the value of  $n$ . Its value is at this point undetermined, but we will argue in the next section that

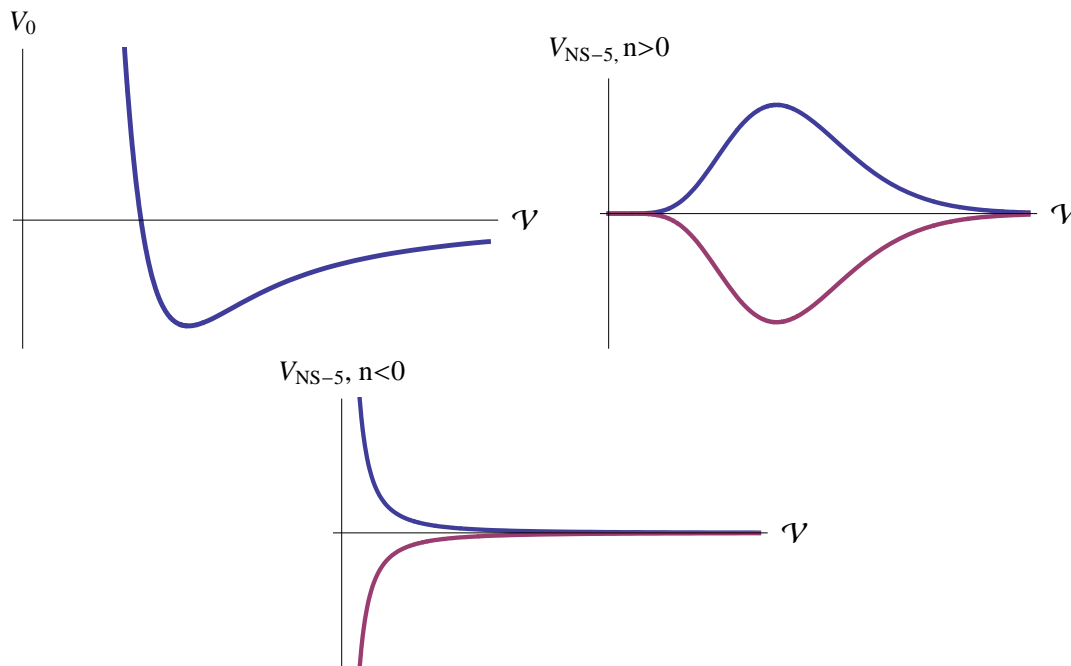
$$n = -3 - \frac{\chi}{12\pi}. \quad (2.19)$$

This is the value obtained in the  $N = 2$  theory, and we assume that its order of magnitude remains the same in the orientifolded  $N = 1$  theory.

For  $n > 0$ , the contribution from the five-brane is positive and will therefore certainly increase the value of the potential at the minimum. Depending on the strength  $\hat{B}$  and the location of the contribution it can produce new vacua or uplift the existing minimum to a de Sitter vacuum. With  $n$  positive and  $\hat{B}$  negative, the five-brane yields a negative energy contribution to the potential function. Depending on the parameters, it will introduce new vacua or lower the potential energy at the location of the existing vacuum. The situation is also interesting if the existing scenarios do not stabilize the volume and we obtain a run-away potential which behaves as  $V_0 \simeq \mathcal{V}^m$ , with  $m$  negative. A five-brane contribution with positive  $n$  can then provide the volume stabilization. The possible signs of  $n$  and  $\hat{B}$  and their results are summarized in the table below.

We have outlined the model and the qualitative behavior, and will now investigate the numerics. The general scenario is described by

$$V = V_0(\mathcal{V}) + V_{\text{NS5}}, \quad V_{\text{NS5}} = \hat{B} \mathcal{V}^n e^{-\mathcal{V}/g_s^2}. \quad (2.20)$$



**Figure 1:** General shape of the three terms in the potential. The first is the AdS minimum of  $V_{\text{pert}}$ . The second and third are the NS5-brane contribution  $\hat{B}\mathcal{V}^n e^{-\mathcal{V}/g_s^2}$ , with  $n$  positive and negative, respectively. The possible signs of  $\hat{B}$  are given by the upper line ( $\hat{B}$  positive) and the lower line ( $\hat{B}$  negative).

$n$	$\hat{B}$	Result
+	+	Increased vacuum energy (uplift), with de Sitter vacua
+	-	Decreased vacuum energy, increased number of minima
-	+	Increased vacuum energy (uplift)
-	-	Decreased vacuum energy

**Table 1:** Possible signs of  $n$  and  $\hat{B}$  and their effect on the vacuum structure.

We have not really specified  $V_0$  here; one can choose a favorite scenario in which the Kähler moduli are stabilized, and then investigate the influence of the NS5-brane contribution. The expression for  $V_{\text{NS5}}$  will, in general, be more complicated: if we take a non-trivial  $W$  (e.g. including D3-brane instantons as in (2.8)), there will be mixing terms consisting of Kähler perturbations and superpotential perturbation in terms like  $|D_i W|^2$ . These terms are typically suppressed, or at most of the same order as in (2.20). In both cases, the numerical analysis below remains the same.

The scalar potential (2.20) contains an overall term  $\exp(K_{\text{CS}}(U, \bar{U}))$ , whose exact value depends on the details of the complex structure moduli stabilization. We set this factor to unity to compare with other models.<sup>4</sup> The KKLT scenario stabilizes the volume at

<sup>4</sup>In  $V_{\text{NS5}}$  it was absorbed in the factor  $\hat{B}$ .

$V_{\text{NS5}}$	$\mathcal{V} = 10$	$\mathcal{V} = 13$	$\mathcal{V} = 20$	$\mathcal{V} = 60$	$\mathcal{V} = 130$
$g_s = 0.5$	$10^{-10}$	$10^{-15}$	$10^{-26}$	$10^{-93}$	$10^{-213}$
$g_s = 1$	$10^2$	$10^0$	$10^{-1}$	$10^{-16}$	$10^{-44}$

**Table 2:** Values of the NS5-brane instantons corrections, for different string couplings and volumes.

$\mathcal{V} = \mathcal{O}(100)$  at a minimum of  $V_0 = -2 \cdot 10^{-15}$ . The LVS on  $\mathbb{P}^4_{[1,1,1,6,9]}$  stabilizes the volume at  $\mathcal{V} = \mathcal{O}(10^{12})$  with a value of  $V_0 = -6 \cdot 10^{-37}$  [9].

Using  $B = -10, W_0 = 1$  (implying  $\hat{B} = 90/8g_s^{-3}$ ) and an Euler number of  $\chi = -300$ , we find (using (2.19)) the values displayed in table 2.

For very weakly coupled strings,  $g_s < 0.5$ , the effect of the NS5-brane instanton is negligible. For  $g_s = 0.5$  one sees from the table that a volume of order  $\mathcal{V} \simeq 15$  yields corrections to the potential that cannot be ignored in a KKLT scenario. For higher values of the string coupling constant,  $g_s = 1$ , the corrections to the potential at  $\mathcal{V} \simeq 60$ , are of the same order as the value of the KKLT-potential at its minimum. In that case, NS5-brane instantons cannot be ignored and can change the KKLT AdS vacuum to become dS, although fine tuning is required.

Quantum corrections (both in  $\alpha'$  and  $g_s$ ) become very important at those scales. The exact form of those corrections is not known, but we can make some rough estimates. The first correction in  $\alpha'$  scales as  $\hat{\xi}^3/\mathcal{V}^4$ , where  $\hat{\xi} = \xi g_s^{-3/2}$ , and higher corrections are expected to be further suppressed by factors  $\hat{\xi}/\mathcal{V}$ . For the values  $g_s = 1, \mathcal{V} = 50, \chi \sim -300$  we find  $\hat{\xi} \sim 0.7$  and

$$V_{\text{NS5}} \sim \frac{\xi^5}{\mathcal{V}^6}.$$

This can be of the same order as next-to-subleading corrections in  $\alpha'$ .

## 2.4 Fivebranes and orientifold projections

We have shown the influence of a NS5-brane instanton on the scalar potential. There is, however, a subtlety in the microscopic string theory that needs to be addressed. The NS5-brane instanton arises from the wrapping of the 10-dimensional NS5-brane soliton solution. In 10 dimensions, the NS5-brane is the magnetic source of the NS-NS  $B_2$  field. Such a wrapping naturally arises when we compactify six internal dimensions. In such a compactification, the 4-dimensional part of the  $B_2$  field is dualized to an axion  $\sigma$ , and the NS5-brane instanton yields exponential corrections of the form

$$e^{-\frac{\mathcal{V}}{g_s} |Q| + iQ\sigma},$$

where  $Q$  is the instanton charge. As usual, the instanton action contains an imaginary part, that distinguishes between instantons ( $Q > 0$ ) and anti-instantons ( $Q < 0$ ). Microscopically, this distinction arises when one has to specify the orientation of the wrapping relative to the orientation of the CY. However, the field  $\sigma$  gets projected out in an orientifold; see e.g. [26]. In the ten-dimensional picture, this corresponds to saying that the space-time part

of  $B_2$  gets projected out, but the NS5-brane couples magnetically to this field. These considerations seem to lead, on the one hand, to the conclusion that an NS5-brane instanton cannot exist, at least not in the traditional sense.

On the other hand, one could argue that at the level of the effective action, all even combinations in  $\sigma$  survive the projection. Examples are  $\exp(-\mathcal{V}/g_s^2)$  and  $\exp(-\mathcal{V}/g_s^2) \cos \sigma$ . The last term can be interpreted as an instanton ( $Q = 1$ ) — anti-instanton pair ( $Q = -1$ ). Such a pair would annihilate, unless some other mechanism stabilizes the pair. They cannot be separated in the internal manifold as e.g. for D3-branes, because they wrap the entire CY. This suggests that they cannot preserve  $N = 1$  SUSY after the orientifold projection, but in the next section, we show that they still can be written in a  $N = 1$  supergravity action. So we are led to the conclusion that instantons do remain present after taking the orientifold projection.

The situation can be described with the following diagram:

$$\begin{array}{ccc}
 \text{II String theory/CY} & \longrightarrow & N = 2, \text{ effective supergravity description} \\
 \downarrow \text{ orientifold} & & \downarrow \text{ orientifold} \\
 \text{II String theory/CY}_{\text{or}} & \longrightarrow & N = 1, \text{ effective supergravity description}
 \end{array} \tag{2.21}$$

Starting from the full-fledged string theory in the top left, one can obtain an effective  $N = 2$  supergravity description, containing effects from NS5-brane instantons. We then orientifold by simply putting  $\sigma = 0$  (and other fields) in this action and obtain an  $N = 1$  theory. This procedure shows that there is a contribution from NS5-branes consistent with  $N = 1$  SUSY. However, one could argue that the correct way to proceed is to incorporate the orientifold projection in string theory, and then calculate the low-energy effects of a NS5-brane. It would be interesting to compare these two approaches; we leave this question open for further investigation.

The situation is better understood before the orientifold projection, when we still have  $N = 2$ . We will now turn to this setting.

### 3. The $N=2$ scenario

In this section we will describe our results in the more stringent language of  $N = 2$  supergravity. In this setting we have good control over the possible quantum corrections. Furthermore, there are no subtleties with the orientifold projection of the NS5-brane, as discussed at the end of the previous section.

Although the previous section dealt with IIB string compactifications, we will change in this section to type IIA models. The reason is of technical origin, as the dimension of the hypermultiplet moduli space in IIA is given by  $4(h_{1,2} + 1)$  (as opposed to  $4(h_{1,1} + 1)$  for IIB), and  $h_{1,2}$  can be set to zero for rigid CY's. This yields a four-dimensional moduli space, which simplifies the analysis. Moreover, NS5-brane instantons were analyzed in these models in [15, 16], and we will make use of these results. We expect that the results for IIA carry over to IIB.

### 3.1 Gauged and ungauged N=2 supergravity

In ungauged  $N = 2$  supergravity, the moduli space has the local product structure

$$\mathcal{M}^K \times \mathcal{M}^Q. \quad (3.1)$$

For type IIA strings, the special Kähler manifold  $\mathcal{M}^K$  has dimension  $2h^{1,1}$  and is spanned by the scalars in the vector multiplets, corresponding to the deformation of the Kähler form. The quaternionic-Kähler space  $\mathcal{M}^Q$  is spanned by the scalars in the hypermultiplets and is  $4(h^{1,2} + 1)$  dimensional. The manifold  $\mathcal{M}^K$  is described in terms of a prepotential  $F(X^I)$ , where  $I = 0, \dots, h^{1,1}$ . In supergravity, this can be any holomorphic function of the  $X^I$  variables of degree two. The prepotentials obtained from IIA string theory have the specific form

$$F(X) = \frac{1}{3!} \frac{d_{ijk} X^i X^j X^k}{X^0} + \frac{i}{2} \zeta(3) \chi(CY) X^0 X^0 - i \sum_{k_a} n_{k_a} \text{Li}_3(e^{2\pi i k_a X^a / X^0}), \quad (3.2)$$

where the first term is a tree-level contribution, the second is a perturbative one-loop correction and the last terms are the nonperturbative worldsheet instanton contributions. Note that there is only one perturbative correction, so the perturbative regime is under complete control. The geometry of  $\mathcal{M}^Q$  is known at tree-level and at one-loop. It is argued in [24] that higher loop corrections can be absorbed into field redefinitions, and if so, the entire perturbative corrected geometry is known [24, 25, 27]. If we restrict ourselves to a rigid CY manifold, which has  $h^{2,1} = 0$  by definition, there is only one hypermultiplet, which is called the universal hypermultiplet (UHM).

Gauged supergravities arise when isometries on the moduli space are gauged. They give rise to scalar potentials that are consistent with  $N = 2$  supersymmetry. Microscopically, gauged supergravities arise when fluxes (in the RR and NS-NS sector) are turned on. For the purpose of our paper, it suffices to look at abelian isometry groups.

The scalar potential is determined by the geometrical data of the moduli space, such as the choice of killing vectors  $k_I$  and their corresponding moment maps  $\vec{\mu}_I$ . For further details on the gauging, we refer to appendix B and references therein. The result for the scalar potential is

$$V = -4 \left[ 2G_{\alpha\beta} k_I^\alpha k_J^\beta + 3\vec{\mu}_I \cdot \vec{\mu}_J \right] \frac{X^I \bar{X}^J}{N_{MN} X^M \bar{X}^N} - 4N_{MN} X^M \bar{X}^N \mathcal{M}_{IJ} N^{IK} N^{JL} \vec{\mu}_K \cdot \vec{\mu}_L. \quad (3.3)$$

In this formula,  $G_{\alpha\beta}$  is the metric on the hypermultiplet space. The gauged isometries are represented by the Killing vectors  $k_I^\alpha$  and their moment maps  $\vec{\mu}_I$ . The matrices  $N_{MN}$  and  $\mathcal{M}_{IJ}$  are defined by

$$N_{IJ} = -iF_{IJ} + i\bar{F}_{IJ},$$

$$\mathcal{M}_{IJ} = \frac{1}{[N_{MN} X^M \bar{X}^N]^2} [N_{IJ} N_{KL} - N_{IK} N_{JL}] \bar{X}^K X^L,$$

where  $F_I = \partial_I F$  etc. In our conventions, both  $G_{\alpha\bar{\beta}}$  and  $\mathcal{M}_{IJ}$  are negative definite, so the first term is a positive contribution. The second is negative, whereas the last one is positive. There is an additional term in non-abelian gaugings which can be omitted for our analysis.

### 3.2 Including NS5-brane corrections

We will now make those expressions explicit for the UHM. The perturbatively corrected metric  $G_{\alpha\beta}$  on the UHM space is given by [25]

$$ds_{\text{UHM}}^2 = \frac{r+2c}{r^2(r+c)} dr^2 + \frac{r+2c}{r^2} (d\chi^2 + d\varphi^2) + \frac{r+c}{r^2(r+2c)} (d\sigma + \chi d\varphi)^2. \quad (3.4)$$

The four bosonic fields are an axion  $\sigma$  from the dualization of the NSNS twoform  $B_2$ , two RR scalars  $\chi, \varphi$  and the four-dimensional dilaton  $g_4$

$$r \equiv e^{\phi_4} = \frac{1}{g_4^2} = \frac{\mathcal{V}}{g_s^2}. \quad (3.5)$$

The relation between the four-dimensional string coupling constant  $g_4$  and the ten-dimensional string coupling constant  $g_s$  is important, as it will introduce factors of the volume into our future expressions.

The constant  $c$  encodes the one-loop correction and is proportional to the Euler number

$$c = -\frac{\chi(\text{CY})}{12\pi} = -\frac{h^{1,1}}{6\pi},$$

where the second equality holds on a rigid CY. The contributions from a single NS5-brane instanton to the UHM metric have been derived in [16, 15]. To leading order in the semiclassical approximation, the metric reads

$$ds_{\text{UHM}}^2 = \frac{r+2c}{r^2(r+c)} dr^2 + \frac{r+2c}{r^2} (1-Y) d\chi^2 + \frac{r+2c}{r^2} (1+Y) d\varphi^2 + \frac{2}{r} \tilde{Y} d\chi d\varphi + \frac{r+c}{r^2(r+2c)} (d\sigma + \chi d\varphi)^2. \quad (3.6)$$

The quantities  $Y$  and  $\tilde{Y}$  are defined as<sup>5</sup>

$$Y = 4C(2\chi^2 - 1)r^{-1-c} \cos(\sigma)e^{-r-\frac{1}{2}\chi^2-c}, \quad \tilde{Y} = 4C(2\chi^2 - 1)r^{-1-c} \sin(\sigma)e^{-r-\frac{1}{2}\chi^2-c}.$$

The factor  $C$  is a numerical constant which could not be determined. This solution has a shift symmetry associated with  $\varphi$  which we can gauge, using the graviphoton as gauge field. The field  $\varphi$  is obtained by expanding the 10-dimensional RR field  $\widehat{C}_3$  over one of the  $2(h^{2,1} + 1) = 2$  cycles in  $H^3$ ; gauging the isometry associated with  $\varphi$  has a microscopic interpretation of adding NS flux over this cycle [28]. Moreover, the shift symmetry is not broken by NS5-brane instantons, so there is no obstruction in gauging this isometry by fluxes in the presence of instantons [29, 30].

---

<sup>5</sup>Compared to [16] we have taken  $\chi_0 = 0$ . Its dependence can easily be restored.

The gauging of this isometry leads to a scalar potential of the type given in (3.3), with a Killing vector  $k = \partial_\varphi$ . Upon inserting the prepotential (3.2) without the worldsheet instantons, one finds that the moment maps drop out of the equation for the potential (3.3), and only the norm of the Killing vector remains. The only dependence on the vector multiplet moduli comes from the factor  $(N_{MN}X^M X^N)^{-1}$ . The details of this calculation can be found in appendix B.

Without the worldsheet instanton corrections, we then find the scalar potential

$$\begin{aligned}
 V &= \frac{2}{4\mathcal{V} + e} \left[ -2G_{\alpha\beta} k^\alpha k^\beta \right] \\
 &= \frac{4}{4\mathcal{V} + e} \left( \frac{4(r + 2c)^2 + 4(r + c)\chi^2}{r^2(r + 2c)} + 16 C e^{-c-r-\chi^2/2} r^{-2-c} (2\chi^2 - 1) \cos(\sigma) \right), \quad (3.7)
 \end{aligned}$$

where  $e = \frac{1}{2}\zeta(3)\chi(CY)$  and  $C$  is the undetermined overall constant. Reinstating all volume factor dependencies using (3.5), this has the schematic form

$$V = V_0 + \tilde{C}\mathcal{V}^{-3-c}e^{-\mathcal{V}/g_s^2},$$

where  $\tilde{C} = 16Cg_s^{4+2c}e^{-c-\chi^2/2}(2\chi^2 - 1)\cos(\sigma)$ . We also neglect the correction due to  $e$  in the 2nd term, because it is subleading.

We see how a NS5-brane contribution can be included into  $N = 2$  type IIA supergravity. It would be interesting to repeat this exercise including the worldsheet instantons.

We now truncate this theory to obtain an  $N = 1$  description.

### 3.3 Truncation to $N = 1$

To clarify the relation to the previous section, we will now perform a truncation of this theory. We make an orientifold inspired truncation to  $N = 1$  at the level of the effective action (see figure (2.21)). A similar truncation has been done in [21]. We follow the orientifold rules from [26]. Because we merely truncate the theory, there should still be a local product structure as in (3.1), but now the product is between two Kähler manifolds. Furthermore, for simplicity we restrict ourself to the cubic prepotential and therefore put  $e = 0$ .

The universal hypermultiplet loses half of its fields under truncation to become a chiral  $N = 1$  multiplet. We keep the four-dimensional dilaton  $r$  and project out the axion  $\sigma$ . From the RR scalars  $\chi, \varphi$  we can choose which we keep. We gauged the isometry on  $\varphi$ , which corresponds to a NS-flux on the cycle of  $\varphi$ . The relevant part of the expansion of  $\hat{H}_3$  and  $\hat{C}_3$  is given by

$$\hat{C}_3 = \chi\alpha + \varphi\beta \quad (3.8)$$

$$\hat{H}_3 = p\alpha + q\beta. \quad (3.9)$$

The field  $\hat{C}_3$  is expanded over a basis of the third cohomology group  $H^3$ , given by three-forms  $\alpha, \beta$ , which give the four-dimensional fields  $\chi, \varphi$ . The flux of  $\hat{B}_2$  is likewise expanded, with flux parameters  $p, q$ . Under an orientifold, the RR form  $\hat{C}_3$  and the NS-NS flux  $\hat{H}_3 = d\hat{B}_2$  are even and odd respectively. We gauge the isometry associated with  $\varphi$ , so we

want to keep the flux parameter  $q$ . This implies that  $\beta$  should be an odd cycle. In the expansion of the even form  $\widehat{C}_3$  we only keep even forms, and hence  $\varphi$  gets projected out.

The metric then truncates to

$$ds_{\text{UHM}}^2 = \frac{r+2c}{r^2(r+c)} dr^2 + \frac{r+2c}{r^2} (1-Y) d\chi^2, \tag{3.10}$$

where we have put  $\sigma = 0$  in  $Y$ . The perturbatively corrected scalar potential is, with  $e = 0$

$$V_0 = \frac{4}{\mathcal{V}} \frac{(r+2c)^2 + (r+c)\chi^2}{r^2(r+2c)}, \tag{3.11}$$

and the NS5-brane instantons yields equation (3.7) with  $\sigma = 0$

$$V_{\text{NS5}} = \frac{16}{\mathcal{V}} C e^{-c-r-\chi^2/2} r^{-2-c} (2\chi^2 - 1), \tag{3.12}$$

where  $C$  is independent of vector multiplet scalars.

We want to express these quantities in terms of a Kähler and superpotential. The Kähler potential  $K$  is a sum of the Kähler potential  $K_k$  for the truncated Kähler moduli and a potential  $K_Q$  for the truncated universal hypermultiplet. In the Kähler sector, we have the Kähler potential [26]

$$K_k = -\ln(\mathcal{V}),$$

which follows from the choice of the cubic prepotential we made in the  $N = 2$  calculation, earlier in this section. The important property of this Kähler potential is its no-scale structure  $K_k^{i\bar{j}} K_i^k K_{\bar{j}}^k = 3$ .

For the scalar potential we use expression (2.6)

$$V = e^K \left( K^{\alpha\bar{\beta}} D_\alpha W D_{\bar{\beta}} \bar{W} - 3|W|^2 \right) = \frac{1}{\mathcal{V}} e^{K_Q} (K_Q^{z\bar{z}} D_z W \bar{D}_{\bar{z}} \bar{W}),$$

where the no-scale structure in directions orthogonal to the truncated universal hypermultiplet has been used.

The perturbative part of the metric (3.10) and the potential (3.11) are now exactly reproduced by the Kähler potential and superpotential [31]

$$\begin{aligned} K_Q &= -2 \ln [(z + \bar{z})^2 - 16c], \\ W &= 16z. \end{aligned}$$

These are formulated in terms of the chiral field  $z$  defined by

$$z = 2\sqrt{r+c} + i\chi.$$

We use conventions for which  $ds^2 = 4K_{z\bar{z}} dz d\bar{z}$ , as in [31].

We now also want to describe the NS5-brane instanton contribution. This scales as

$$\exp\left(-r - \frac{\chi^2}{2} - c\right) = \exp\left(\frac{1}{16}(z^2 - 6z\bar{z} + \bar{z}^2)\right),$$



which is not holomorphic in  $z$ . Therefore, we cannot correct the superpotential with such a term. The correction will take place in the Kähler potential and in the definition of the  $N = 1$  chiral field. Both the metric (3.10) and the potential (3.7) are reproduced up to leading order by the chiral field and the Kähler potential

$$\begin{aligned}
 z &= 2\sqrt{r+c} + i\chi + 2Ce^{-r-\frac{1}{2}\chi^2-c}r^{-2-c}(\sqrt{r}(1-2\chi^2) + i\chi(2\chi^2-5)), \quad (3.13) \\
 K_Q &= -2\ln[(z+\bar{z})^2 - 16c] + C\exp\left[\frac{1}{16}(z^2 - 6z\bar{z} + \bar{z}^2)\right] \times \\
 &\quad \times 4^{9+2c}\frac{(z+\bar{z})^{-4-2c}(1+\frac{1}{2}(z-\bar{z})^2)}{(z-3\bar{z})(\bar{z}-3z)}, \\
 W &= 16z.
 \end{aligned}$$

Because our four-dimensional dilaton  $r$  contains a factor of the volume  $\mathcal{V}$ , the leading term in the Kähler potential is equal to

$$K_Q = K_0 + B\mathcal{V}^n e^{-\mathcal{V}/g_s^2}, \quad (3.14)$$

where we have defined

$$B = 64Ce^{-c-\chi^2/2}(1-2\chi^2)g_s^{-2n}, \quad n = -3 - c. \quad (3.15)$$

The overall factor  $B$  can depend on other moduli. In this setting, the leading dependence on  $\chi$  is explicit in (3.15). We expect (3.14) to hold also for non-rigid CY with  $h_{1,2} \neq 0$ . In that case the factor  $B$  presumably depends on the other hypermultiplet scalars.

This confirms our proposal of (2.12) and (2.19), where we apply it to type IIB string theory.

## Acknowledgments

We would like to thank Lilia Anguelova, Tim Baarslag, Ralph Blumenhagen, Jan de Boer and Frank Saueressig for discussions and correspondence. This work is partially supported by the European Union RTN network MRTN-CT-2004-005104 and INTAS contract 03-51-6346.

## A. Calculations in $N = 1$

In this appendix we give some details of the calculations which have been used in section 2.

We want to calculate the derivatives and inverses for the Kähler potentials (2.4)

$$K_0 = -\ln(-i(\tau - \bar{\tau})) - 2\ln\left(\mathcal{V} + \frac{\hat{\xi}}{2}\right), \quad (A.1)$$

$$K_{NS5} = B\mathcal{V}^n \exp(-\mathcal{V}\tau_2^2), \quad (A.2)$$

where we use

$$\begin{aligned}\tau &= l + ie^\phi, \quad \tau_2 = \text{Im}\tau, \\ 6\mathcal{V} &= d_{ijk}t^i t^j t^k = d_{ij}t^i t^j = d_i t^i, \\ \hat{\xi} &= \xi(\text{Im}\tau)^{3/2}.\end{aligned}$$

We introduce  $d^{ij}$  as the inverse of  $d_{ij}$ , and denote  $\mathcal{A} := \mathcal{V} + \hat{\xi}/2$ . We first consider the case where  $G^a = 0$ . From  $T_i = \tau_i + ib_i$  (equation (2.1) for  $G^a = 0$ ) we find

$$\frac{\partial t^j}{\partial T_i} = \frac{1}{4}d^{ij}, \quad \frac{\partial \mathcal{V}}{\partial T_i} = \frac{1}{8}t^i.$$

We can then calculate

$$\begin{aligned}K_{T_i}^0 &= -\frac{1}{4}\frac{t^i}{\mathcal{A}}, \quad K_\tau^0 = i\tau_2^{-1}\left(\frac{3\hat{\xi}}{4\mathcal{A}} + \frac{1}{2}\right), \\ K_{T_i \bar{T}_j}^0 &= \frac{G^{ij}}{\mathcal{A}^2}, \quad G^{ij} = -\frac{1}{16}\mathcal{A}d^{ij} + \frac{1}{32}t^i t^j, \\ K_{i\bar{\tau}}^0 &= \frac{3i}{32\mathcal{A}^2}\tau_2^{-1}\hat{\xi}t^i, \\ K_{\tau\bar{\tau}}^0 &= \frac{1}{16}\tau_2^{-2}\mathcal{A}^{-2}(4\mathcal{V}^2 + \mathcal{V}\hat{\xi} + 4\hat{\xi}^2).\end{aligned}$$

This can be inverted to give

$$\begin{aligned}K_0^{\tau\bar{\tau}} &= \frac{4\mathcal{V} - \hat{\xi}}{\mathcal{V} - \hat{\xi}}\tau_2^2, \\ K_0^{i\bar{\tau}} &= -\frac{3i\hat{\xi}}{\mathcal{V} - \hat{\xi}}\tau_2 d_i, \\ K_0^{i\bar{j}} &= -8(2\mathcal{V} + \hat{\xi})d_{ij} + \frac{4\mathcal{V} - \hat{\xi}}{\mathcal{V} - \hat{\xi}}d_i d_j.\end{aligned}$$

We then find a familiar result, which leads directly to (2.10):

$$K_0^{\alpha\bar{\beta}} K_\alpha^0 K_\beta^0 = 3 + \frac{3\hat{\xi}(\mathcal{V}^2 + 7\mathcal{V}\hat{\xi} + \hat{\xi}^2)}{(\mathcal{V} - \hat{\xi})(2\mathcal{V} + \hat{\xi})^2}. \tag{A.3}$$

For the NS5-brane contribution we obtain (we write  $K_5 = K^5 = K_{\text{NS5}}$ )

$$\begin{aligned}K_{\tau\bar{\tau}}^5 &= -\frac{1}{4}K_5\mathcal{V}((\tau - \bar{\tau})^2\mathcal{V} + 2), \\ K_{i\bar{\tau}}^5 &= \frac{i}{8}B \exp(-\mathcal{V}\tau_2^2)\mathcal{V}^n(-\mathcal{V}\tau_2 + n + 1)\tau_2 t^i, \\ K_{i\bar{j}}^5 &= \frac{1}{64}B t^i t^j \mathcal{V}^{n-2}(\mathcal{V}^2\tau_2^4 - \mathcal{V}2n\tau_2^2 + (n-1)n) \exp(-\mathcal{V}\tau_2^2).\end{aligned}$$

We are interested in the leading term in the potential, so we want to investigate the powers of the volume. If we denote volume powers with  $[\cdot]$ , then

$$[\mathcal{V}] = 1, \quad [d_{ijk}] = 0, \quad [t^i] = \frac{1}{3}, \quad [d_{ij}] = \frac{1}{3}.$$

All the one-instanton terms are multiplied by  $\exp(-\mathcal{V}\tau_2^2)$ . To determine the leading term, we have to find the highest power of the volume  $\mathcal{V}$  in the polynomial which appears in front of this exponent. Therefore, we do not include the factor  $\exp(-\mathcal{V}\tau_2^2)$  in the counting, or equivalently we put  $[\exp(-\mathcal{V}\tau_2^2)] = 0$ .

The various terms have the following leading volume dependencies:

$$\begin{aligned} [K_0^{i\bar{j}}] &= 4/3 & [K_{i\bar{j}}^5] &= n + 2/3 & [K_0^{i\bar{j}} K_{\bar{j}k}^5 K_0^{\bar{k}l}] &= n + 10/3 \\ [K_0^{i\bar{\tau}}] &= -1/3 & [K_{i\bar{\tau}}^5] &= n + 4/3 & [K_0^{i\bar{\tau}} K_{\bar{\tau}k}^5 K_0^{\bar{k}l}] &= n + 7/3 \\ [K_0^{\tau\bar{\tau}}] &= 0 & [K_{\tau\bar{\tau}}^5] &= n + 2 & [K_0^{i\bar{\tau}} K_{\bar{\tau}\tau}^5 K_0^{\bar{k}l}] &= n + 4/3. \end{aligned}$$

The leading contribution is given by (we use  $\simeq$  here to denote equality up to subleading terms)

$$K_5^{i\bar{j}} = K_0^{i\bar{\alpha}} K_{\bar{\alpha}\beta}^5 K_0^{\beta\bar{j}} \simeq K_0^{i\bar{l}} K_{l\bar{k}}^5 K_0^{k\bar{j}} \simeq B\mathcal{V}^{n+2} g_s^{-4} \exp(-\mathcal{V}/g_s^2) d_i d_j.$$

In the potential we find then  $[K_5^{i\bar{j}} |\partial_i K_0|^2] = (n + 10/3) - 2/3 - 2/3 = n + 2$ . The other term is  $[K_0^{i\bar{j}} \partial_i K_0^5 \partial_{\bar{j}} K_0] = 4/3 + (n + 1/3) - 2/3 = n + 1$ , which is subleading with respect to the terms above. Then the leading contribution to the scalar potential is given by

$$\begin{aligned} V &\simeq -e^{K_0} K_{\text{NS5}}^{i\bar{j}} |\partial_i K_0|^2 |W_0|^2 & (A.4) \\ &= -\frac{9}{8} B\mathcal{V}^n g_s^{-3} |W_0|^2 \exp(-\mathcal{V}/g_s^2). \end{aligned}$$

Let us now consider the effects of non-zero  $G^a$ , to clarify the statements made after equation (2.15). From the definition (2.1)

$$T_i = \tau_i + ib_i + \frac{i}{\tau - \bar{\tau}} d_{iab} G^a (G - \bar{G})^c,$$

we find that

$$\frac{\partial t^j}{\partial T_i} = \frac{1}{4} d^{ij}, \quad \frac{\partial t^i}{\partial G^a} = \frac{1}{4} d^{ij} d_{jab} b^b, \quad \frac{\partial t^i}{\partial \tau} = \frac{i}{2} d_{iab} b^a b^b,$$

and hence

$$\frac{\partial \mathcal{V}}{\partial T_i} = \frac{1}{8} t^i, \quad \frac{\partial \mathcal{V}}{\partial G^a} = \frac{1}{2} t^j d_{jac} b^c = \frac{1}{2} d_{ac} b^c, \quad \frac{\partial \mathcal{V}}{\partial \tau} = \frac{i}{4} t^j d_{jab} b^a b^b = \frac{i}{4} d_{ab} b^a b^b.$$

In the last two expressions the factor  $t^j$  is bound with the factor  $d_{jac}$  and cannot combine with a  $d_j$  to form a power of the volume.

The expression for (2.15) also contains inverse metrics. If we use the expressions for the tree-level Kähler metric in [19], we can explicitly determine the volume dependence, and we find

$$\begin{aligned} K_0^{ia} K_{ab}^5 K_0^{bj} t^i t^j &\sim \mathcal{V}^{n+2} \\ K_0^{ik} K_{kl}^5 K_0^{lj} t^i t^j &\sim \mathcal{V}^{n+4} \\ K_0^{ia} K_{ak}^5 K_0^{kj} t^i t^j &\sim \mathcal{V}^{n+3}, \end{aligned}$$

and the leading term does not contain the fields  $G^a$ . We do not know if this property holds when we include quantum corrections to the Kähler potential, but as quantum corrections are expected to be subleading in the volume, we expect this to be the case.

## B. Potentials in $N = 2$

In this appendix we give some more details of the calculation in the  $N = 2$  setting. This appendix derives a general form of the scalar potential. The next appendix specializes this to the UHM and the Przanowski metric.

We use the formalism from [32] with the vector prepotential

$$F = \frac{1}{3!} \frac{d_{ijk} X^i X^j X^k}{X^0} + \frac{i}{2} \zeta(3) \chi(CY) X^0 X^0. \quad (\text{B.1})$$

From the prepotential we define

$$N_{IJ} = -iF_{IJ} + i\bar{F}_{IJ} = 2\text{Im} F_{IJ},$$

$$\mathcal{M}_{IJ} = \frac{1}{[N_{MN} X^M \bar{X}^N]^2} [N_{IJ} N_{KL} - N_{IK} N_{JL}] \bar{X}^K X^L,$$

and then the scalar potential is given by

$$V = -4g^2 \left[ 2G_{\alpha\beta} k_I^\alpha k_J^\beta + 3\vec{\mu}_I \cdot \vec{\mu}_J \right] \frac{X^I \bar{X}^J}{N_{MN} X^M \bar{X}^N}$$

$$- g^2 N_{MN} X^M \bar{X}^N \mathcal{M}_{IJ} \left[ 4N^{IK} N^{JL} \vec{\mu}_L \cdot \vec{\mu}_L - \frac{f_{KL}^I X^K \bar{X}^L}{N_{PQ} X^P \bar{X}^Q} \frac{f_{MN}^J X^M \bar{X}^N}{N_{PQ} X^P \bar{X}^Q} \right],$$

where  $g$  is an overall factor to make the terms which are a result of the gauging more explicit in the Lagrangian; we put  $g = 1$  from now on. In general, each vector field in the vector multiplets can be used to gauge one of the  $h^{1,1} + 1$  different killing vectors  $k^I$ . In our setting, there is only one isometry  $k$ , which we gauge by the graviphoton. The index  $I$  therefore only attains the value 0. If we use  $\vec{\mu}_I = \delta_I^0 \vec{\mu}$  we obtain

$$V = -\frac{4}{N_{MN} X^M \bar{X}^N} \left( \left[ 2G_{\alpha\beta} k^\alpha k^\beta + 3\vec{\mu}^2 \right] X^0 \bar{X}^0 + (N_{KL} N^{00} - \delta_K^0 \delta_L^0) \bar{X}^K X^L \vec{\mu}^2 \right),$$

and the term depending on  $f_{KL}^I$  is zero for abelian gaugings. We now use the prepotential

$$F = \frac{1}{3!} d_{ijk} \frac{X^i X^j X^k}{X^0} + \frac{1}{2} e X^0 X^0,$$

where  $e = \frac{i}{2} \zeta(3) \chi(CY)$  is purely imaginary. Using  $X^i/X^0 = z^i = b^i + it^i$ , we find

$$F_{00} = \frac{1}{3} d_{ijk} z^i z^j z^k + e \quad \text{Im} F_{00} = d_{ijk} b^i b^j t^k - \frac{1}{3} d_{ijk} t^i t^j t^k$$

$$F_{0i} = -\frac{1}{2} d_{ijk} z^j z^k \quad \text{Im} F_{0i} = -d_{ijk} b^j t^k$$

$$F_{ij} = d_{ijk} z^k \quad \text{Im} F_{ij} = d_{ijk} t^k.$$

Using the abbreviations  $d_{ij} = d_{ijk} t^k$ ,  $d_i = d_{ij} t^j$ ,  $6\mathcal{V} = d_{ijk} t^i t^j t^k$  we find

$$N_{00} = 2d_{ij} b^i b^j - 4\mathcal{V}' \quad N^{00} = \frac{-1}{4\mathcal{V}'}$$

$$N_{0i} = -2d_{ij} b^j \quad N^{0i} = \frac{-b^i}{4\mathcal{V}'}$$

$$N_{ij} = 2d_{ij} \quad N^{ij} = \frac{-b^i b^j}{4\mathcal{V}'} + \frac{1}{2} d^{ij}$$

$$e^{-K} \equiv N_{IJ} X^I \bar{X}^J = (8\mathcal{V} + 2e) X^0 \bar{X}^0.$$
(B.2)

where we have written  $\mathcal{V}' := \mathcal{V} - \frac{1}{2}e$ . For the scalar potential we then finally find equation (3.7)

$$\begin{aligned} V &= -\frac{4}{8\mathcal{V} + 2e} \left( \left[ 2G_{\alpha\bar{\beta}} k^\alpha k^\beta + 3\vec{\mu}^2 \right] + \left( 8\mathcal{V}' \frac{-1}{4\mathcal{V}'} - 1 \right) \vec{\mu}^2 \right), \\ &= \frac{2}{4\mathcal{V} + e} \left[ -2G_{\alpha\bar{\beta}} k^\alpha k^\beta \right]. \end{aligned} \quad (\text{B.3})$$

The moment maps drop out and the scalar potential is positive definite.

## C. The Przanowski metric

In this appendix we repeat some of the results of [33, 15], which are used to determine the NS5-brane one-instanton corrected  $N = 2$  moduli space in section 3.2.

In [33], it has been shown that a four-dimensional quaternionic-Kähler manifold  $M$  can be described in terms of a partial differential equation for a single, real function. Locally, the metric takes the form

$$\begin{aligned} g &= g_{\alpha\bar{\beta}} (dz^\alpha \otimes dz^{\bar{\beta}} + dz^{\bar{\beta}} \otimes dz^\alpha) \\ &= g_{1\bar{1}} dz^1 dz^{\bar{1}} + g_{1\bar{2}} dz^1 dz^{\bar{2}} + g_{2\bar{1}} dz^2 dz^{\bar{1}} + g_{2\bar{2}} dz^2 dz^{\bar{2}} + c.c., \end{aligned} \quad (\text{C.1})$$

where indices  $\alpha, \beta, \bar{\alpha}, \bar{\beta} = 1, 2$ , and we have used the usual convention of complex conjugation  $z^{\bar{\alpha}} := \overline{z^\alpha}$ . The Hermiticity of this metric is encoded in the requirement  $\overline{g_{\alpha\bar{\beta}}} = g_{\beta\bar{\alpha}}$ . The elements  $g_{\alpha\bar{\beta}}$  are now defined in terms of a real function  $h = h(z^\alpha, z^{\bar{\alpha}})$  via

$$g_{\alpha\bar{\beta}} = 2 \left( h_{\alpha\bar{\beta}} + 2\delta_\alpha^2 \delta_{\bar{\beta}}^2 e^h \right), \quad (\text{C.2})$$

where the subscript  $\alpha$  on  $h_\alpha$  indicates differentiation of the function with respect to  $z^\alpha$ . We have changed the sign of our defining function  $h$  with respect to the original function  $u$  used by Przanowski, as it offers a slightly more convenient form to work with.

The differential equation which determines the function  $h$  is the non-linear partial differential equation

$$h_{1\bar{1}} h_{2\bar{2}} - h_{1\bar{2}} h_{\bar{1}2} + (2h_{1\bar{1}} - h_1 h_{\bar{1}}) e^h = 0. \quad (\text{C.3})$$

### C.1 Solutions to the master equation

The equation (C.3) is a difficult partial differential equation. There have been various approaches in the literature which found exact and approximate solutions to the master equation. By imposing additional symmetries on the manifold  $M$ , one can simplify the master equation. Imposing one isometry reduces this equation to the Toda equation [33]. Upon imposing two commuting isometries one obtains the Calderbank-Pedersen metrics [34].

In [15], solutions to the master equation were obtained which corresponded to NS5-brane instantons. The relation between the complex coordinates and the real coordinates is given by

$$z^1 = \frac{1}{2}(u + i\sigma), \quad z^2 = \frac{1}{2}(\chi + i\varphi), \quad u \equiv r - \frac{1}{2}\chi^2 + c \log(r + c).$$

The leading term of the one-instanton contribution is captured by

$$\begin{aligned}
 h &= h_0 + \Lambda, & h_0 &= \log(r + c) - 2 \log r \\
 \Lambda &= Cr^{-2-c} \cos(\sigma) \exp \left[ -r + \frac{1}{2} \chi^2 \right].
 \end{aligned}
 \tag{C.4}$$

From the metric we only need the length of the Killing vector  $k = \partial_\varphi$ , which can be found from (C.1), (C.2) and (C.4) and is given by

$$-G_{\alpha\beta} k^\alpha k^\beta = \frac{4((r + 2c)^2 + (r + c)\chi^2)}{r^2(r + 2c)} + 16Cr^{-2-c}(2\chi^2 - 1) \exp(-c - r - \chi^2/2).$$

Inserting this into (B.3) yields (3.7).

### C.2 Moment maps

Although the moment maps are not present in the scalar potential, we include their calculation for completeness. We follow the conventions on quaternionic-Kähler geometry from [35].

We want to find vielbeins  $a, b$  for the metric (C.1) such that

$$a \otimes \bar{a} + b \otimes \bar{b} + c.c. = ds^2.$$

Using the Ansatz  $a = \alpha dz^1 + \beta dz^2, b = \gamma dz^1 + \delta dz^2$  we find

$$\begin{aligned}
 a &= \sqrt{2h_{1\bar{1}}} dz^1 + \sqrt{2} \frac{h_{1\bar{2}}}{\sqrt{h_{1\bar{1}}}} dz^2, \\
 b &= \sqrt{2} e^{h/2} \sqrt{\frac{h_1 h_{\bar{1}}}{h_{1\bar{1}}}} dz^2.
 \end{aligned}$$

From those, we determine the SU(2) connection one-forms

$$\begin{aligned}
 \omega^1 &= i \frac{e^{h/2}}{\sqrt{h_1 h_{\bar{1}}}} (h_{\bar{1}} dz^2 - h_1 dz^{\bar{2}}), \\
 \omega^2 &= -\frac{e^{h/2}}{\sqrt{h_1 h_{\bar{1}}}} (h_{\bar{1}} dz^2 + h_1 dz^{\bar{2}}), \\
 \omega^3 &= -\frac{i}{2} \left( h_1 - \frac{h_{1\bar{1}}}{h_{\bar{1}}} + \frac{h_{11}}{h_1} \right) dz^1 \\
 &\quad - \frac{i}{2} \left( h_2 - \frac{h_{1\bar{2}}}{h_{\bar{1}}} + \frac{h_{12}}{h_1} \right) dz^2 + c.c.
 \end{aligned}$$

As a non-trivial check, we can use the tree-level UHM metric, and these one-forms agree with the those obtained in [35]. Notice that the situation drastically simplifies when there is an additional killing vector in the direction  $i(\partial_1 - \partial_{\bar{1}})$ , because then  $h_1 = h_{\bar{1}}$ .

We now gauge the isometry associated with  $\varphi$ . In the complex coordinates, this is the vector

$$k = \frac{1}{2} i (\partial_2 - \partial_{\bar{2}}),$$

where the normalization is such that  $k = \partial_\varphi$ . Calculations of the moment maps is now straight-forward and after some algebra we find

$$\vec{\mu} = \begin{pmatrix} \frac{e^{h/2}}{\sqrt{h_1 h_{\bar{1}}}}(h_1 + h_{\bar{1}}) \\ -i \frac{e^{h/2}}{\sqrt{h_1 h_{\bar{1}}}}(h_1 - h_{\bar{1}}) \\ -h_2 \end{pmatrix},$$

which are real ( $h_2 = h_{\bar{2}}$ ).

The square of the moment maps therefore reads

$$\vec{\mu}^2 = (4e^h + h_2^2) = 4e^h + (\partial_\chi h)^2,$$

where we have used  $\partial_\varphi h = 0$ . This last expression is valid in the coordinates  $(u, \sigma, \chi, \varphi)$ . Changing to the coordinates  $(r, \sigma, \chi, \varphi)$  amounts to changing the derivatives according to

$$\partial_\chi \rightarrow \partial_\chi + \chi \frac{r+c}{r+2c} \partial_r.$$

## References

- [1] M. Graña, *Flux compactifications in string theory: a comprehensive review*, *Phys. Rept.* **423** (2006) 91 [[hep-th/0509003](#)].
- [2] M.R. Douglas and S. Kachru, *Flux compactification*, *Rev. Mod. Phys.* **79** (2007) 733 [[hep-th/0610102](#)].
- [3] R. Blumenhagen, B. Körs, D. Lüst and S. Stieberger, *Four-dimensional string compactifications with D-branes, orientifolds and fluxes*, *Phys. Rept.* **445** (2007) 1 [[hep-th/0610327](#)].
- [4] S. Kachru, R. Kallosh, A. Linde and S.P. Trivedi, *De Sitter vacua in string theory*, *Phys. Rev. D* **68** (2003) 046005 [[hep-th/0301240](#)].
- [5] V. Balasubramanian, P. Berglund, J.P. Conlon and F. Quevedo, *Systematics of moduli stabilisation in Calabi-Yau flux compactifications*, *JHEP* **03** (2005) 007 [[hep-th/0502058](#)].
- [6] J.P. Conlon, F. Quevedo and K. Suruliz, *Large-volume flux compactifications: moduli spectrum and D3/D7 soft supersymmetry breaking*, *JHEP* **08** (2005) 007 [[hep-th/0505076](#)].
- [7] S. Kachru et al., *Towards inflation in string theory*, *JCAP* **10** (2003) 013 [[hep-th/0308055](#)].
- [8] J.P. Conlon and F. Quevedo, *Kähler moduli inflation*, *JHEP* **01** (2006) 146 [[hep-th/0509012](#)].
- [9] M. Berg, M. Haack and E. Pajer, *Jumping through loops: on soft terms from large volume compactifications*, *JHEP* **09** (2007) 031 [[arXiv:0704.0737](#)].
- [10] M. Berg, M. Haack and B. Körs, *Loop corrections to volume moduli and inflation in string theory*, *Phys. Rev. D* **71** (2005) 026005 [[hep-th/0404087](#)].
- [11] O. DeWolfe, A. Giriyavets, S. Kachru and W. Taylor, *Type IIA moduli stabilization*, *JHEP* **07** (2005) 066 [[hep-th/0505160](#)].
- [12] M. Berg, M. Haack and B. Körs, *On volume stabilization by quantum corrections*, *Phys. Rev. Lett.* **96** (2006) 021601 [[hep-th/0508171](#)].

- [13] M. Cicoli, J.P. Conlon and F. Quevedo, *Systematics of string loop corrections in type IIB Calabi-Yau flux compactifications*, *JHEP* **01** (2008) 052 [[arXiv:0708.1873](#)].
- [14] S. Kachru and A.-K. Kashani-Poor, *Moduli potentials in type IIA compactifications with RR and NS flux*, *JHEP* **03** (2005) 066 [[hep-th/0411279](#)].
- [15] S. Alexandrov, F. Saueressig and S. Vandoren, *Membrane and fivebrane instantons from quaternionic geometry*, *JHEP* **09** (2006) 040 [[hep-th/0606259](#)].
- [16] M. Davidse, U. Theis and S. Vandoren, *Fivebrane instanton corrections to the universal hypermultiplet*, *Nucl. Phys. B* **697** (2004) 48 [*Erratum ibid.* **750** (2006) 108] [[hep-th/0404147](#)].
- [17] M. de Vroome and S. Vandoren, *Supergravity description of spacetime instantons*, *Class. and Quant. Grav.* **24** (2007) 509 [[hep-th/0607055](#)].
- [18] M. Gutperle and M. Spalinski, *Supergravity instantons for  $N = 2$  hypermultiplets*, *Nucl. Phys. B* **598** (2001) 509 [[hep-th/0010192](#)].
- [19] T.W. Grimm and J. Louis, *The effective action of  $N = 1$  Calabi-Yau orientifolds*, *Nucl. Phys. B* **699** (2004) 387 [[hep-th/0403067](#)].
- [20] K. Becker, M. Becker, M. Haack and J. Louis, *Supersymmetry breaking and  $\alpha'$ -corrections to flux induced potentials*, *JHEP* **06** (2002) 060 [[hep-th/0204254](#)].
- [21] V. Balasubramanian and P. Berglund, *Stringy corrections to Kähler potentials, SUSY breaking and the cosmological constant problem*, *JHEP* **11** (2004) 085 [[hep-th/0408054](#)].
- [22] S.S. AbdusSalam, J.P. Conlon, F. Quevedo and K. Suruliz, *Scanning the Landscape of Flux Compactifications: vacuum Structure and Soft Supersymmetry Breaking*, *JHEP* **12** (2007) 036 [[arXiv:0709.0221](#)].
- [23] K. Becker, M. Becker and A. Strominger, *Five-branes, membranes and nonperturbative string theory*, *Nucl. Phys. B* **456** (1995) 130 [[hep-th/9507158](#)].
- [24] D. Robles-Llana, F. Saueressig and S. Vandoren, *String loop corrected hypermultiplet moduli spaces*, *JHEP* **03** (2006) 081 [[hep-th/0602164](#)].
- [25] I. Antoniadis, R. Minasian, S. Theisen and P. Vanhove, *String loop corrections to the universal hypermultiplet*, *Class. and Quant. Grav.* **20** (2003) 5079 [[hep-th/0307268](#)].
- [26] T.W. Grimm and J. Louis, *The effective action of type IIA Calabi-Yau orientifolds*, *Nucl. Phys. B* **718** (2005) 153 [[hep-th/0412277](#)].
- [27] L. Anguelova, M. Roček and S. Vandoren, *Quantum corrections to the universal hypermultiplet and superspace*, *Phys. Rev. D* **70** (2004) 066001 [[hep-th/0402132](#)].
- [28] J. Louis and A. Micu, *Type II theories compactified on Calabi-Yau threefolds in the presence of background fluxes*, *Nucl. Phys. B* **635** (2002) 395 [[hep-th/0202168](#)].
- [29] A.-K. Kashani-Poor and A. Tomasiello, *A stringy test of flux-induced isometry gauging*, *Nucl. Phys. B* **728** (2005) 135 [[hep-th/0505208](#)].
- [30] L. Anguelova and K. Zoubos, *Five-brane instantons vs flux-induced gauging of isometries*, *JHEP* **10** (2006) 071 [[hep-th/0606271](#)].
- [31] F. Saueressig, U. Theis and S. Vandoren, *On de Sitter vacua in type IIA orientifold compactifications*, *Phys. Lett. B* **633** (2006) 125 [[hep-th/0506181](#)].



- [32] B. de Wit, M. Roček and S. Vandoren, *Gauging isometries on hyperKähler cones and quaternion- Kähler manifolds*, *Phys. Lett. B* **511** (2001) 302 [[hep-th/0104215](#)].
- [33] M. Przanowski, *Killing vector fields in self-dual, euclidean Einstein spaces with  $\Lambda \neq 0$* , *J. Math. Phys.* **32** (1991) 1004.
- [34] D.M.J. Calderbank and H. Pedersen, *Selfdual Einstein metrics with torus symmetry*, *J. Diff. Geom.* **60** (2002) 485 [[math.DG/0105263](#)].
- [35] M. Davidse, F. Saueressig, U. Theis and S. Vandoren, *Membrane instantons and de Sitter vacua*, *JHEP* **09** (2005) 065 [[hep-th/0506097](#)].